20. Astronomers use the following relationship to determine the distance, d, to a star.

$$F = \frac{L}{4\pi d^2}$$

For a particular star the following measurements are recorded:

apparent brightness, $F = 4.4 \times 10^{-10} \,\mathrm{W \, m^{-2}}$

luminosity, $L = 6.1 \times 10^{30} \,\mathrm{W}$

Based on this information, the distance to this star is

- A $3.3 \times 10^{19} \, \text{m}$
- B 1.5×10^{21} m
- C $3.7 \times 10^{36} \, \text{m}$
- D $1.1 \times 10^{39} \, \text{m}$
- E 3.9×10^{39} m.